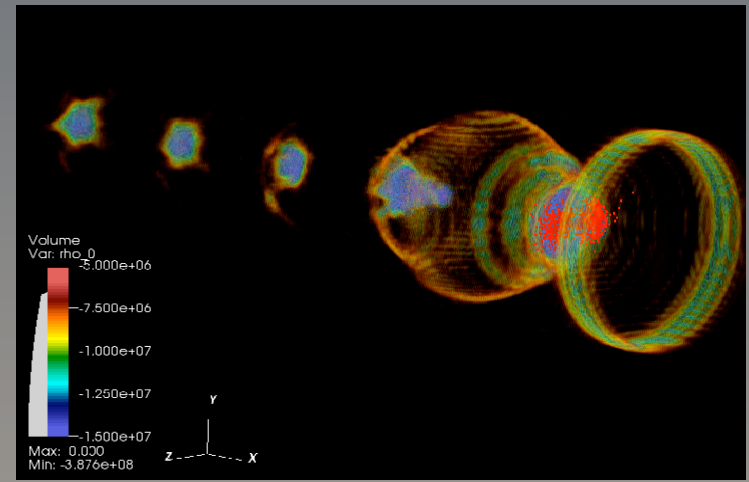
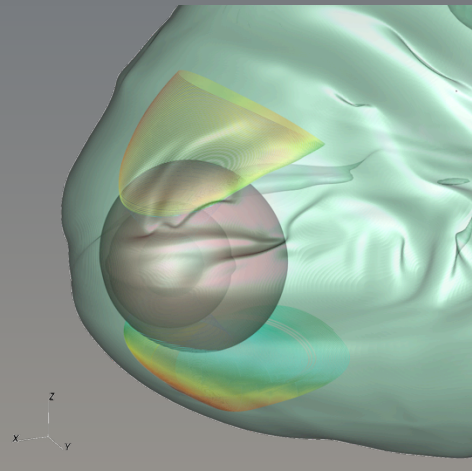
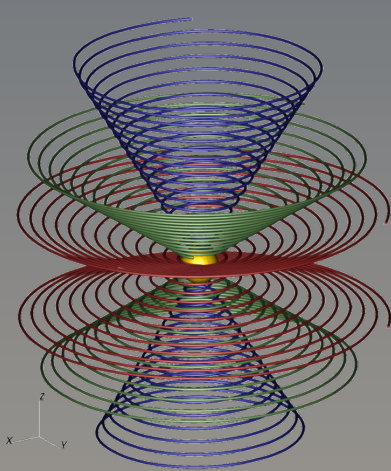


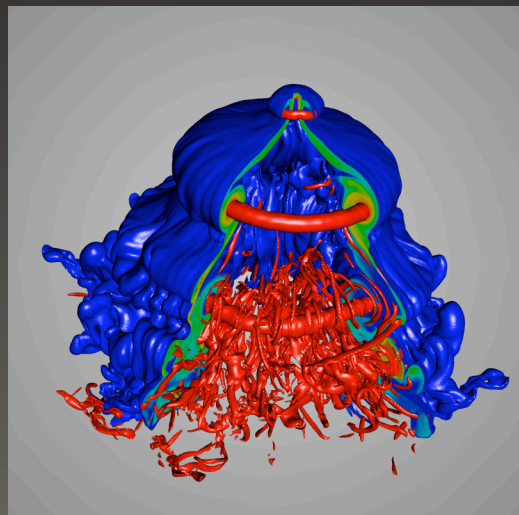
Recent Advances in VisIt: AMR Streamlines and Query-driven Visualization

Gunther H. Weber

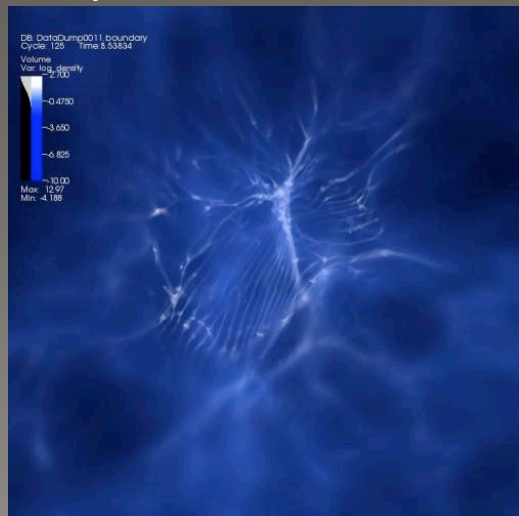
Lawrence Berkeley National Laboratory / UC Davis



- Richly featured visualization and analysis tool for large data sets
- Built for five use cases:
 - Data exploration
 - Visual debugging
 - Quantitative analysis
 - Presentation graphics
 - Comparative analysis
- Data-parallel client server model, distribution on per patch-basis
- Use of meta-data / contracts to reduce amount of processed data



[Argon bubble subjected to shock Jeff Greenbough, LLNL]

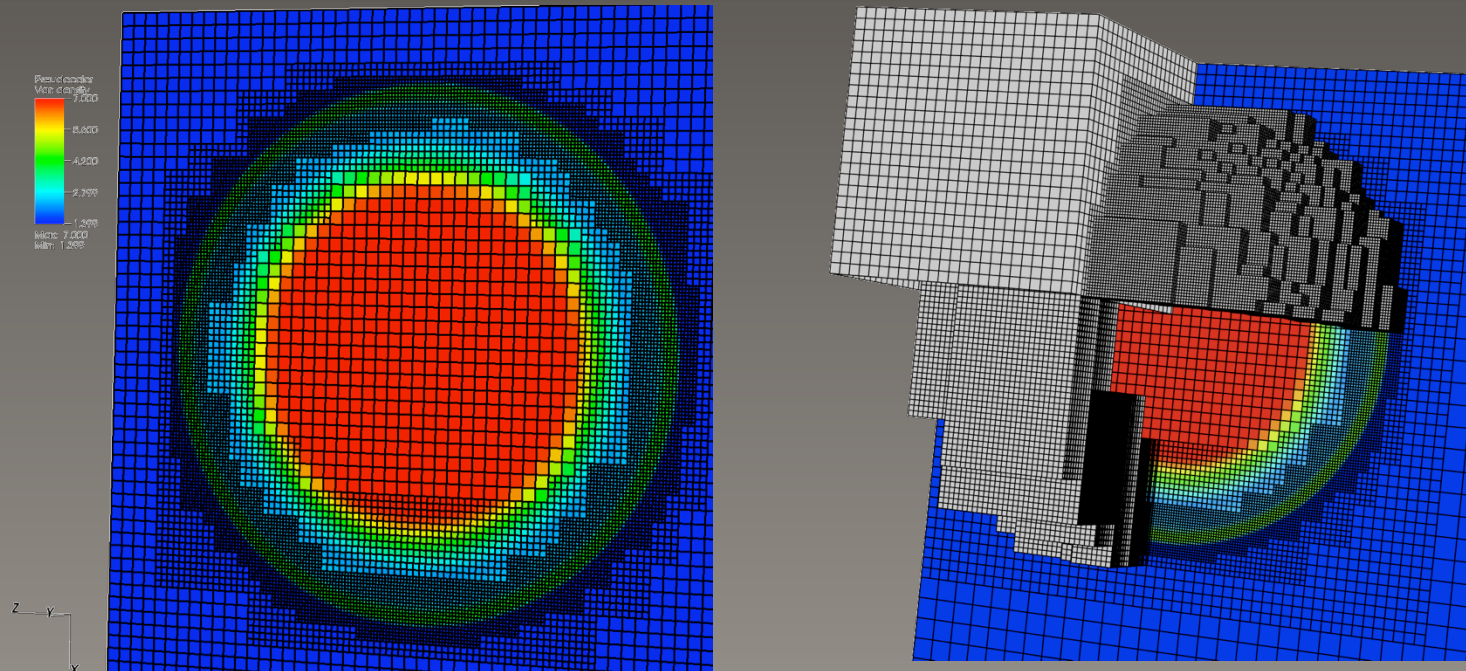


[Logarithm of gas/dust density in Enzo star/galaxy simulation, Tom Abel & Matthew Turk, Kavli Institute]

➔ Ideal basis for specialized AMR visualization tool replacement

VisIt and AMR Data

- Supported as “first-class” data type
- Handled via “ghost-cells”: Coarse cells that are refined are marked “ghost” in the lower level
- Work on rectilinear grids and skip ghost cells or “remove” results produced in ghost cells later on



Streamlines for Adaptive Mesh Refinement Data

Joint work with

Eduard Deines¹, Christoph Garth¹, Brian Van Straalen²,
Sergey Borovikov³, Daniel F. Martin² and Ken I. Joy^{1,2}

¹ Institute for Data Analysis and Visualization, University of California, Davis

² Computational Research Division, Lawrence Berkeley National Laboratory

³ Department of Physics, University of Alabama, Huntsville



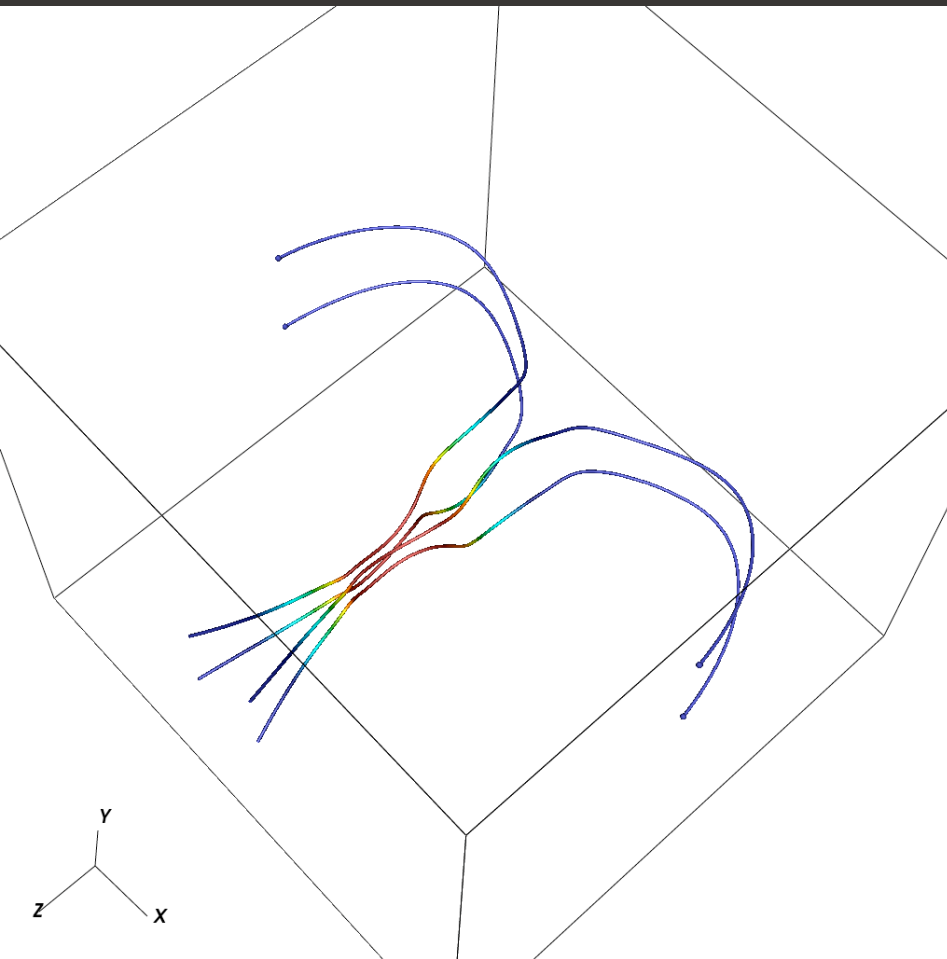
VACET



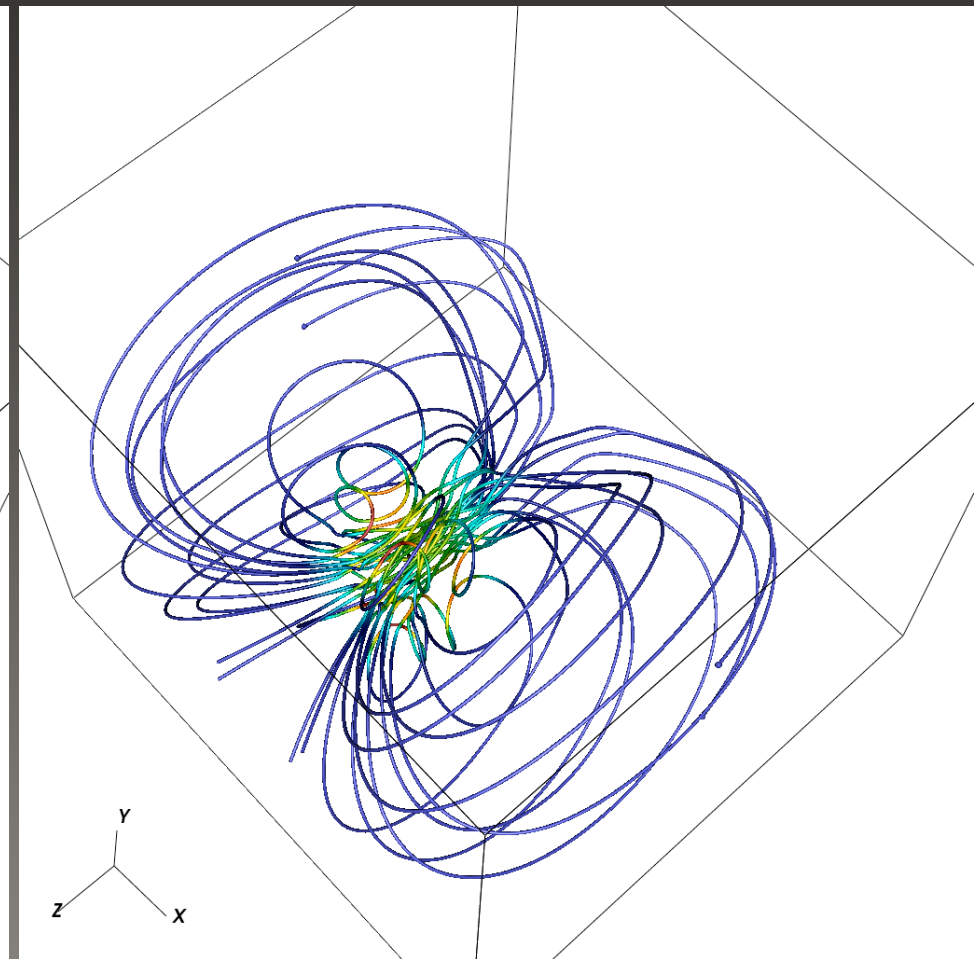
Motivation

- Visit streamlines one year ago:
 - Streamlines cannot cross boundaries between multiple grids of multi-block data
- More recently:
 - Fully parallel implementation that “hands off” streamlines between grids [Pugmire et al, 2009]
 - Publicly available in Visit 1.11 (enhancements in upcoming 1.12 release)
- However:
 - Streamlines still not “AMR” aware (refined cells)

Motivating Example – Vortex Core Merger



Stay in Level 0



“Descend” into AMR hierarchy

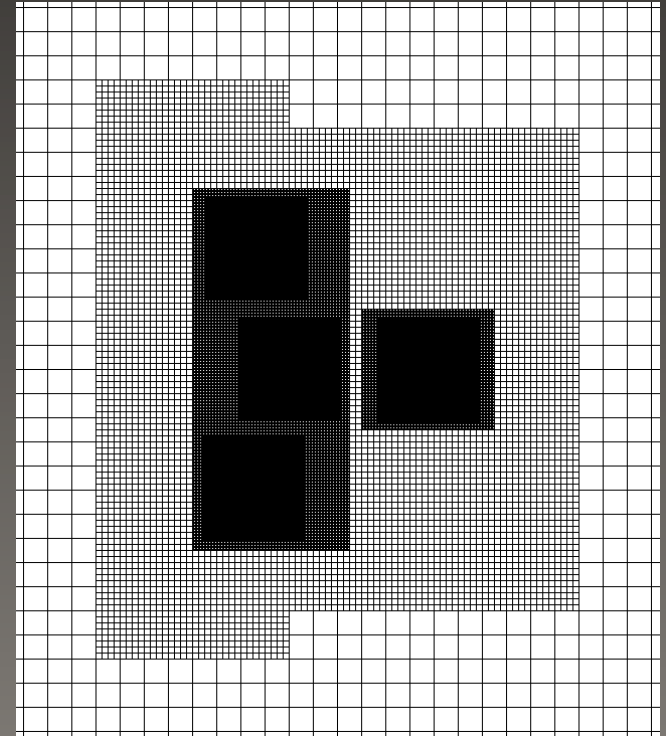
Integral Curves and Streamlines – Definition

- Integral curves in visualization
 - Streamlines
 - Streaklines
 - Pathlines, etc.
- Essential visualization tool providing intuitive understanding of flow data
- Obtained Numerical solution of ordinary differential equation

$$S(t)' = F(S(t)) \text{ or } S(t)' = F(S(t),t)$$

Challenges Posed by Block-structured AMR data sets

- Multiple refinement levels
- Several domains represented as rectilinear grids
- Data in finer levels replaces data in coarser levels
- Cell-centered data



Proper Handling of AMR Hierarchies (1/2)

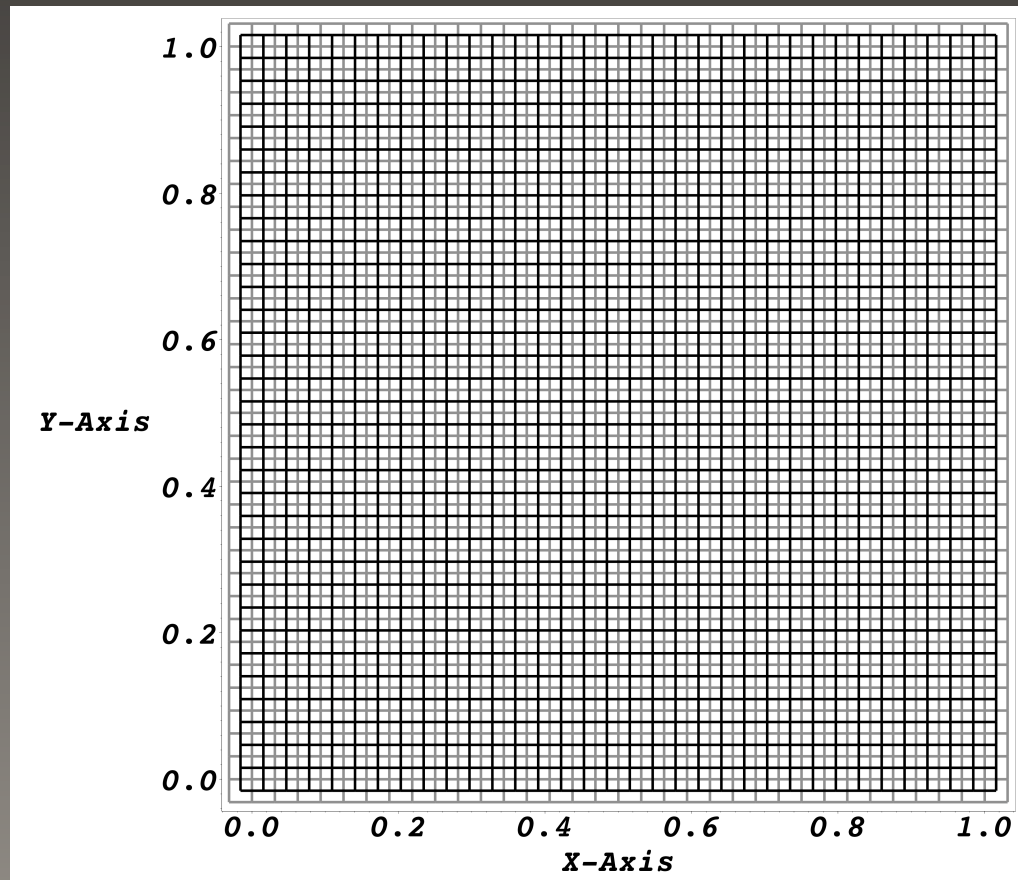
1. Evaluate vector field at each integration step in finest available level:
 - Discontinuity handling by the integration method
 - Higher number of integration steps
 - Higher error after passing the discontinuity (level boundary)
 - All domains needed during the integration

Proper Handling of AMR Hierarchies (2/2)

2. Stop at domain boundaries and restart computation in next domain:
 - Reduces error
 - Reduces number of integration steps
 - Process curve integration in each domain separately (parallel processing possible)
 - Intersection point calculation with domain bounding box using Newton method

Proper Handling of Cell-centered Data

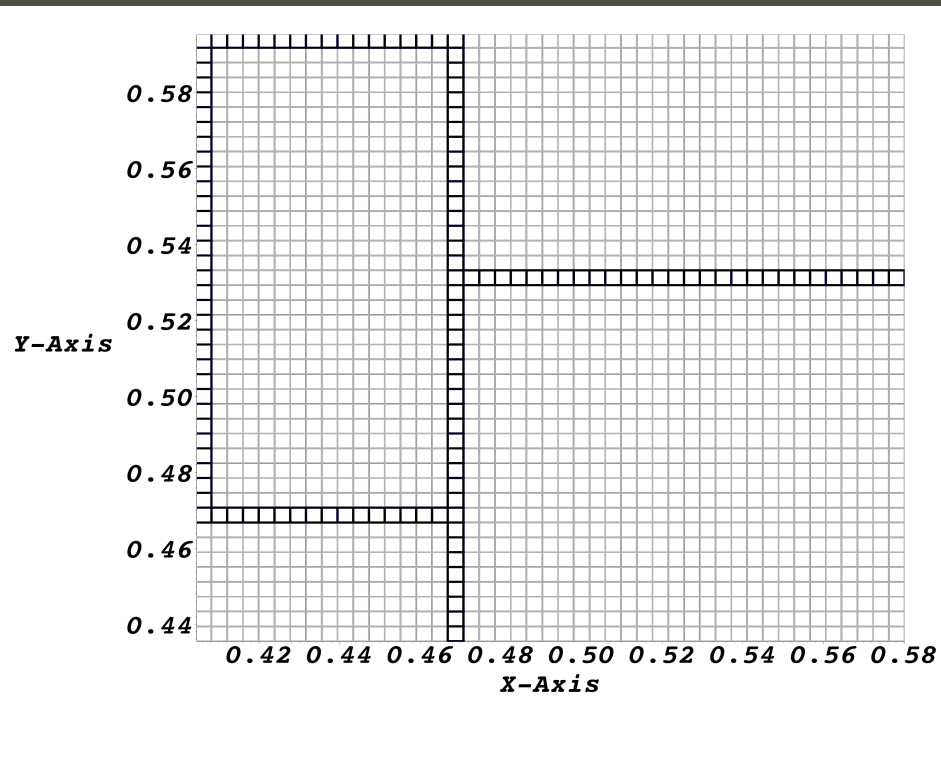
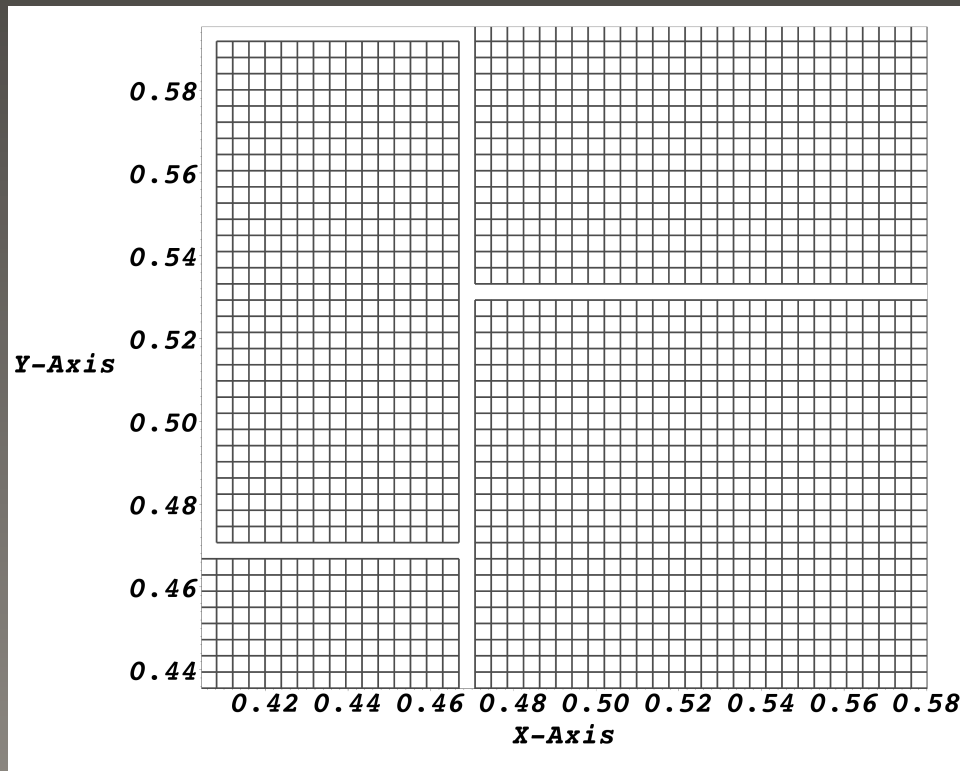
Dual-mesh representation



Proper handling of cell-centered data

“Gaps” between domains

Dual grid using “ghost” cells

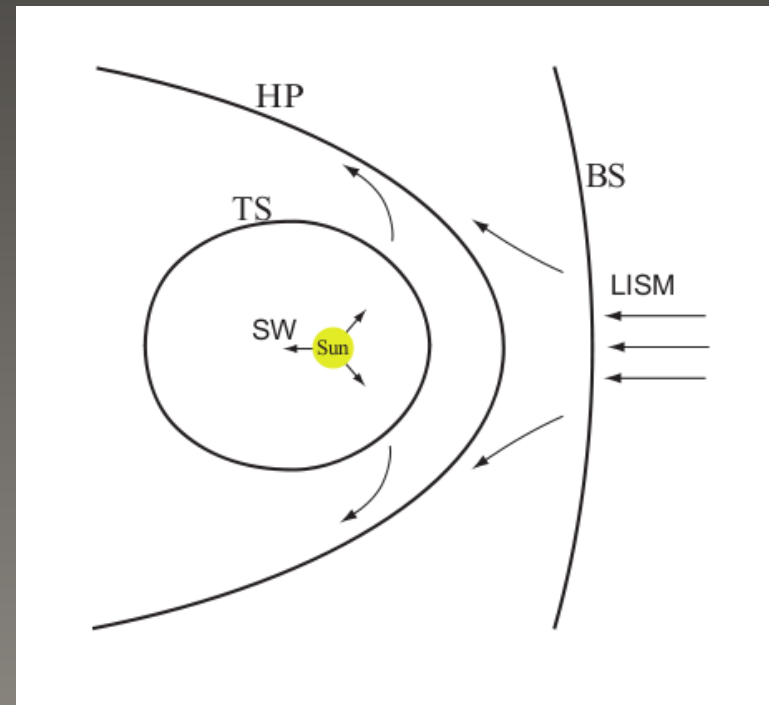


Algorithm

- Start in domain in finest possible level
- Build dual mesh
- Advance integration step
- If step inside nested domains (finer level)
 - Intersect with the bounding box of the finer domain
 - Restart the algorithm inside the finer domain
- If outside domain
 - Intersect with the domain bounding box
 - Restart in the next domain

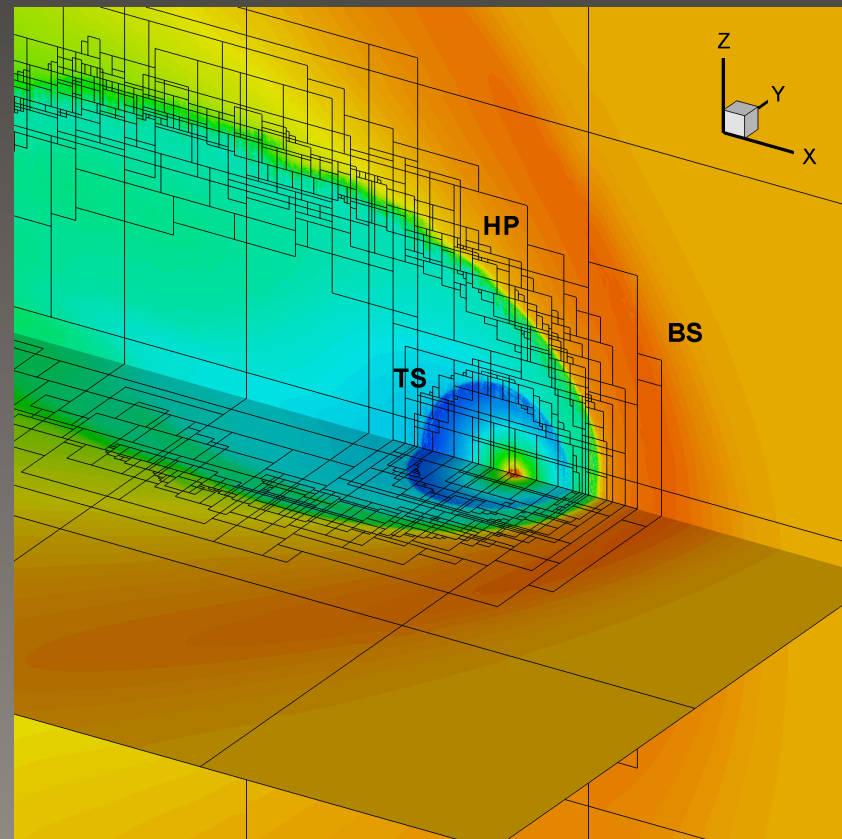
Solar System Simulation

- Interaction of solar wind with interstellar medium
- Termination shock
 - First boundary of solar system
 - End of strong sphere of influence of solar wind
- Heliopause
 - Collision of solar wind with local interstellar medium
 - End of direct sphere of influence of solar wind

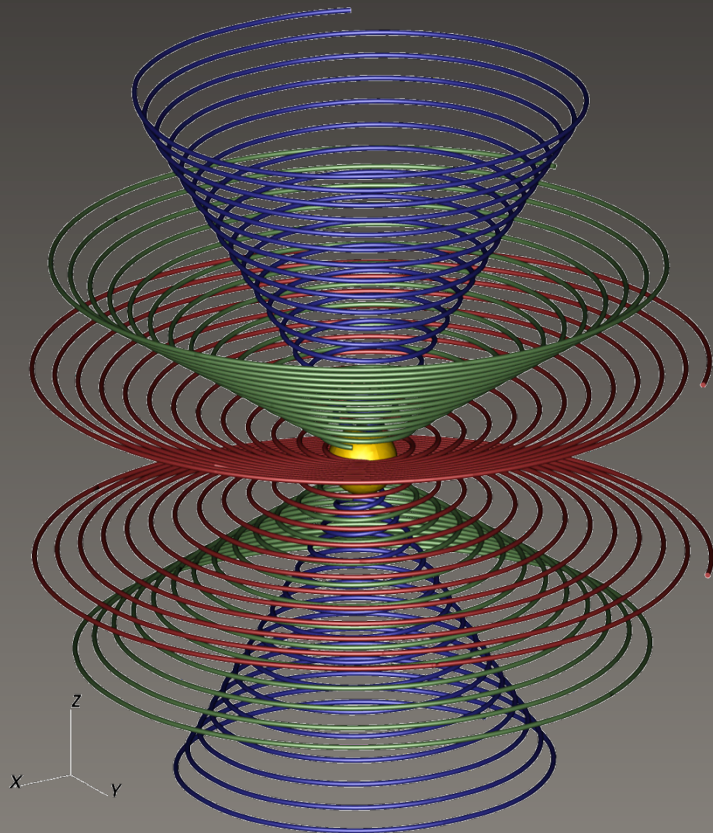


Solar System Simulation

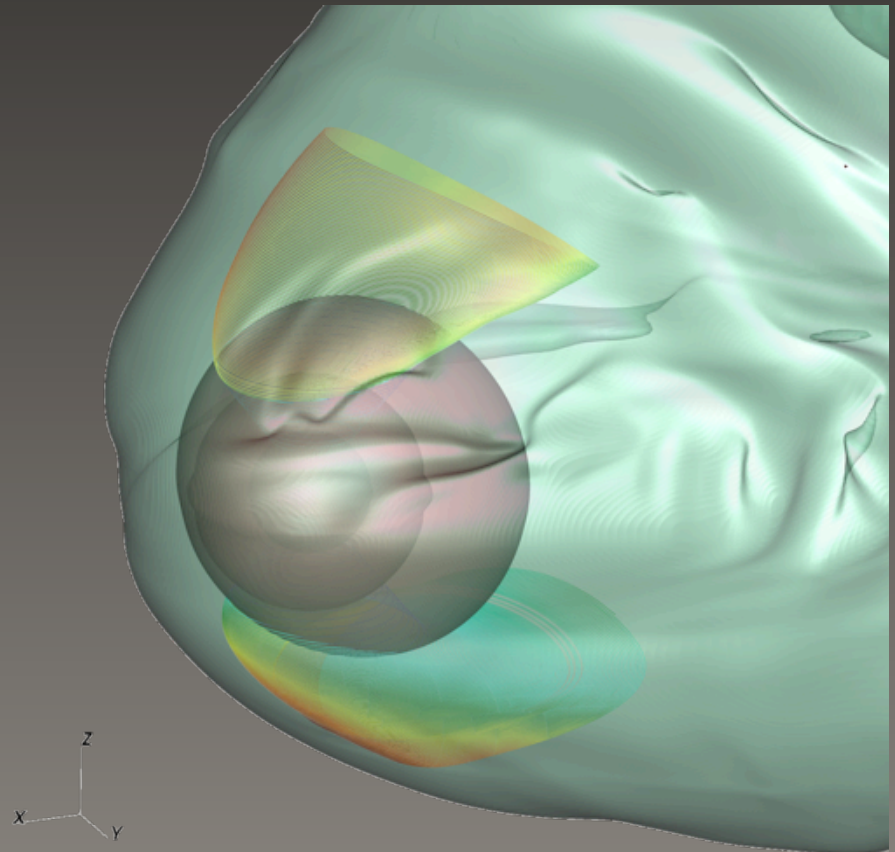
- Computational region about 1000 AU
- Plasma fluctuations 0.01 AU
- Too fine to be modeled without AMR
- AMR Mesh
 - Five refinement levels
 - 20037 domains



Interplanetary Magnetic Field Lines



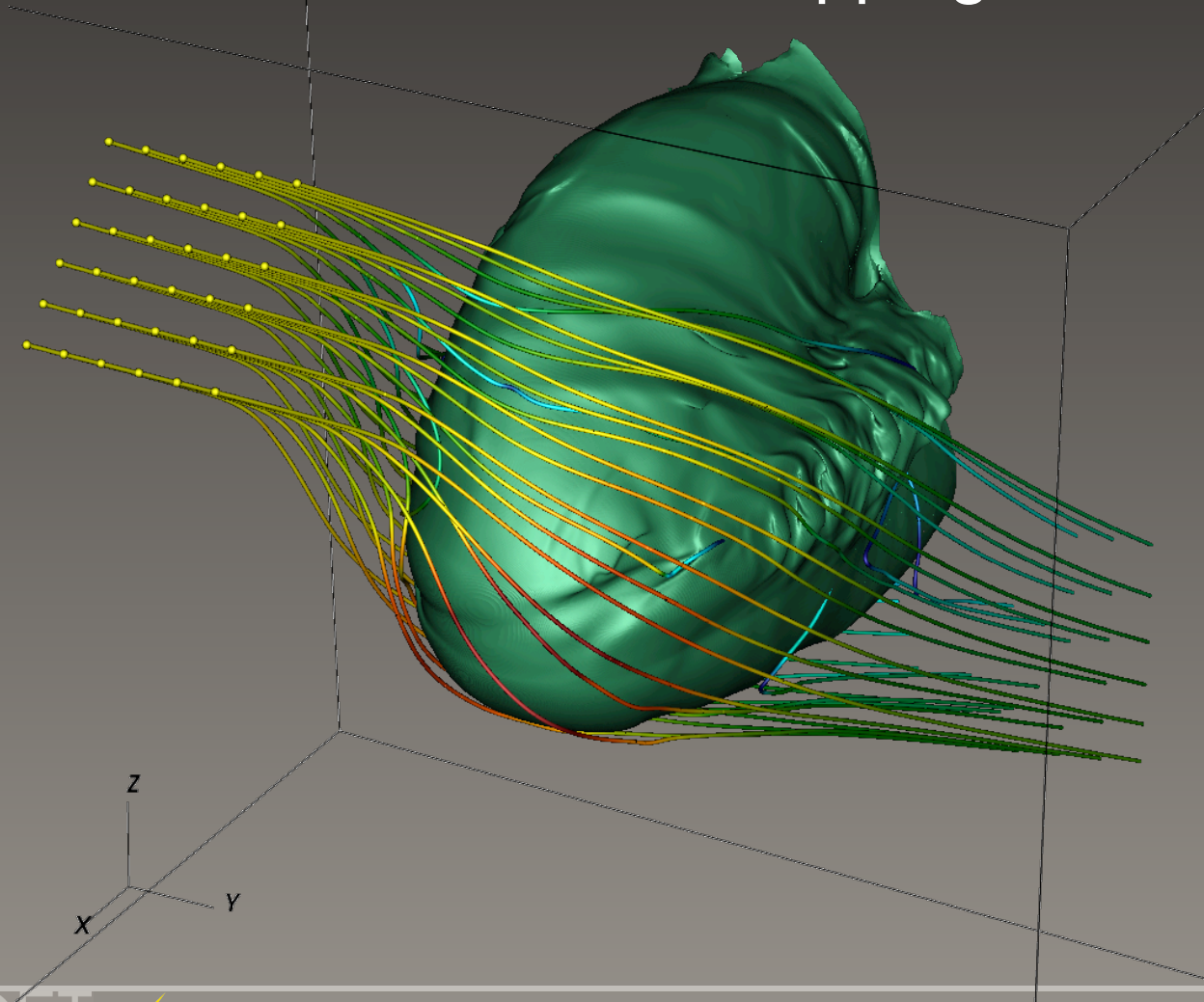
Parker Spiral



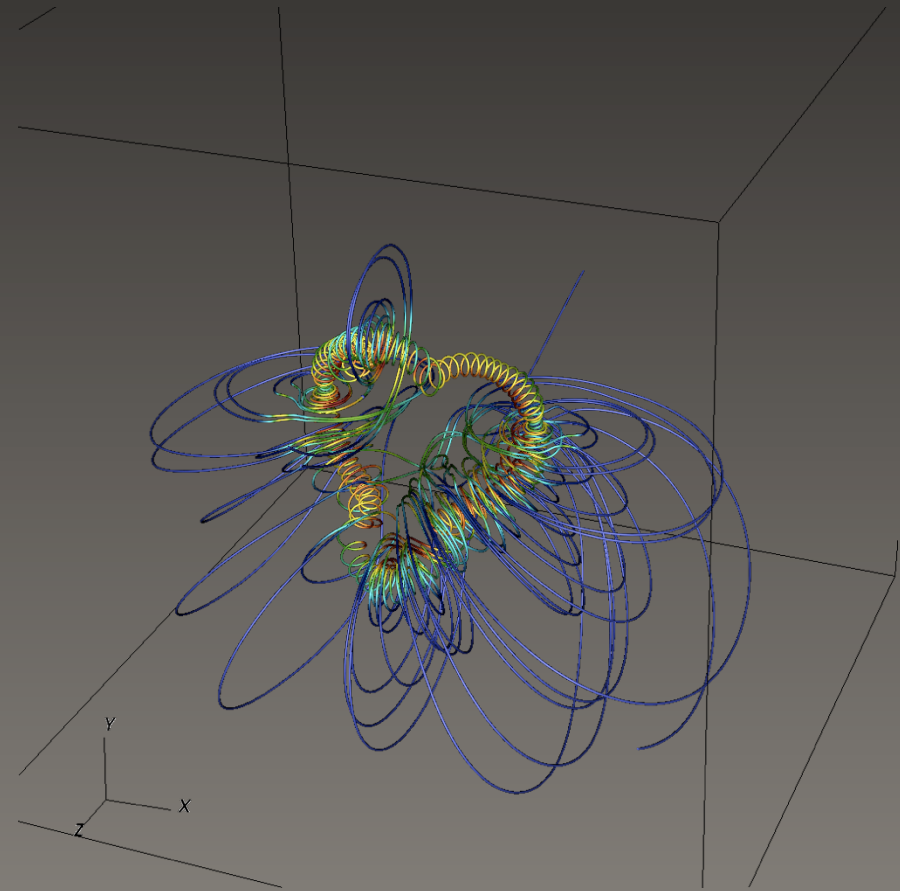
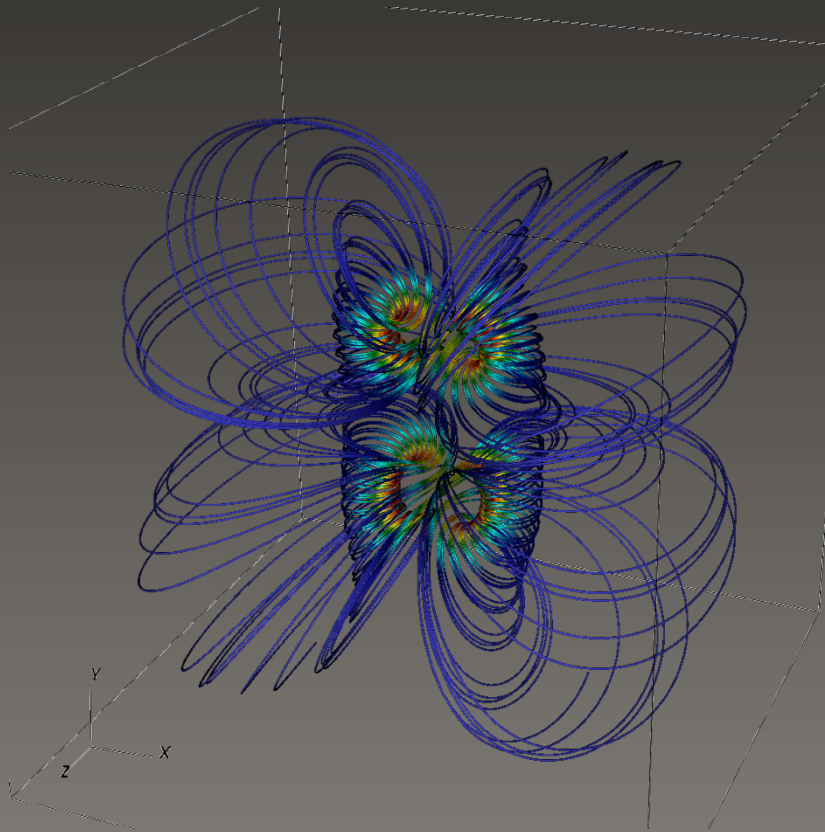
Past the Termination Shock

Interstellar Magnetic Field Lines

Wrapping the Heliopause



Two Incompressible Viscous Vortex Cores



Two distinct vortex rings Merged into a single flow structure

Future work

- Distribute in future VisIt version
- Proper handling of:
 - Embedded boundaries
 - Mapped grids
- Parallelize
- Pathlines, etc.

Analysis of Large-Scale Laser Wakefield Particle Acceleration Simulation

Joint work with

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Jeremy Meredith⁴, Cameron G. R. Geddes⁵,
Estelle Cormier-Michel⁵, Sean Ahern⁴, Peter Messmer⁶,
Hans Hagen², Bernd Hamann^{3,2,1} and E. Wes Bethel^{1,3}

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² Internatl. Research Training Group 1131, TU Kaiserslautern, Germany

³ Institute for Data Analysis and Visualization, University of California, Davis

⁴ Oak Ridge National Laboratory

⁵ LOASIS program of Lawrence Berkeley National Laboratory

⁶ Tech-X Corporation



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Motivation: Laser Wakefield Particle Acceleration



[Image courtesy of <http://worldwakesurfingchampionships.com>]

- **Advantages:**
 - Can achieve electric fields thousands of times stronger than in conventional accelerators
 - ➔ Can achieve high acceleration over very short distance.

Laser Wakefield Particle Acceleration

Simulation

- Performed over 2D and 3D domains using the VORPAL code
- Simulations restricted to window covering only a plasma subset in x direction in beam vicinity
- Simulation window moves along local x axis
- Produces particle and field data (at typically 40-100 timesteps)

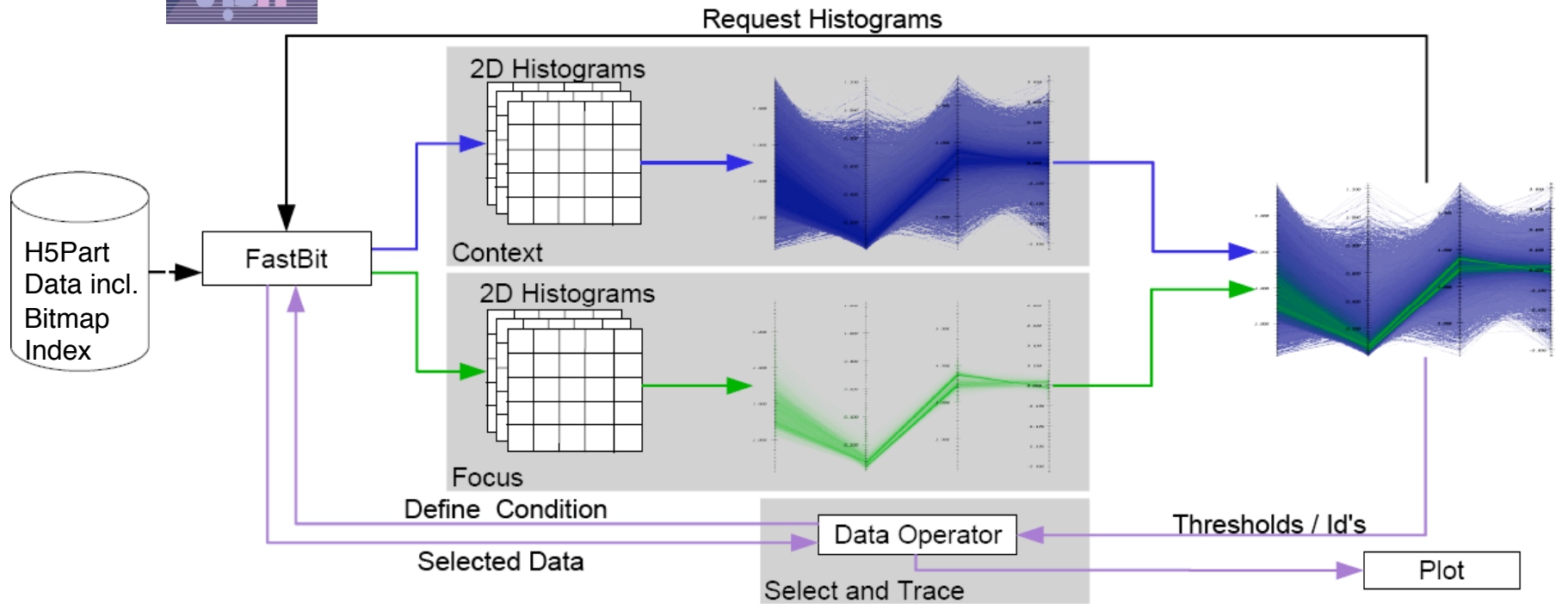
Particle data

- Scattered data with particle location, momentum and identifier
~ $0.4 \cdot 10^6$ – $30 \cdot 10^6$ (in 2D) and ~ $80 \cdot 10^6$ – **$200 \cdot 10^6$** (in 3D) per time step
→ Total size: ~1.5GB – >30GB (in 2D) and ~100GB – **>1TB** (in 3D)

Field data

- Electric field, magnetic field, and RhoJ (regular grid)
 - Resolution: Typically ~0.02-0.03 μm longitudinally, and ~ 0.1-0.2 μm transversely
 - Total size: ~3.5GB - >70GB (in 2D) and ~200GB - >2TB (in 3D)

System Design



References:

- Visit is available from <https://wci.llnl.gov/codes/visit/>
- FastBit is available from <https://codeforge.lbl.gov/projects/fastbit>
- H5Part is available from <http://h5part.web.psi.ch/> or <http://vis.lbl.gov/Research/AcceleratorSAPP/>

Data Selection via FastBit

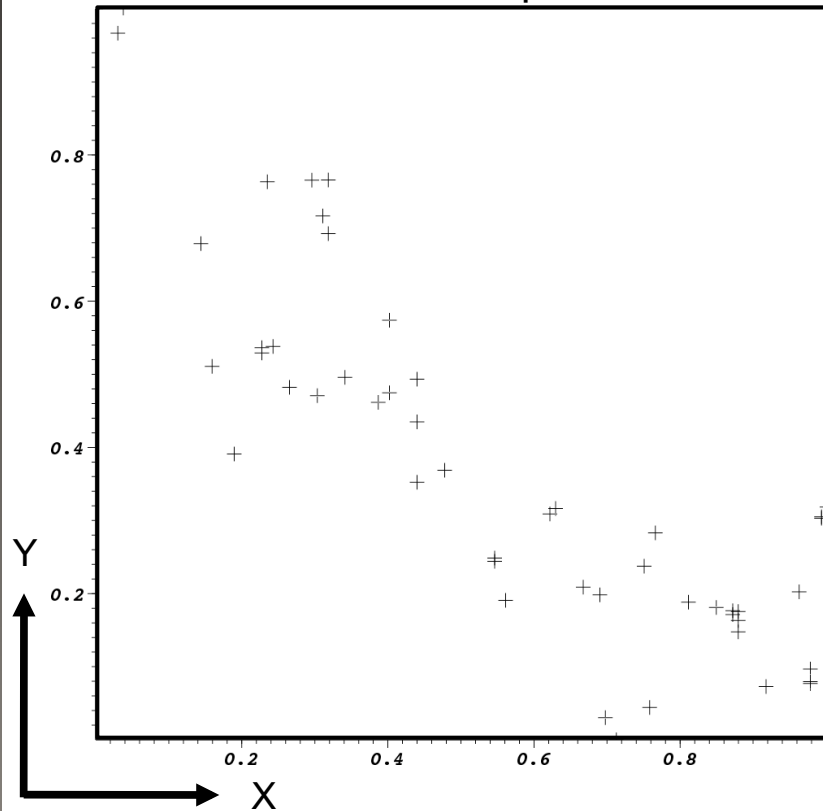
Value	b_0	b_1	b_2	b_3	b_4	b_5
0	1	0	0	0	0	0
1	0	1	0	0	0	0
5	0	0	0	0	0	1
3	0	0	0	1	0	0
1	0	1	0	0	0	0
2	0	0	1	0	0	0
4	0	0	0	0	1	0
...						
	=0	=1	=2	=3	=4	=5

- Use FastBit to accelerate:
 - Computation of conditional histograms for parallel coordinate rendering
 - Multi-dimensional threshold queries for particle of interest identification
 - ID-queries for tracing of particles over time:

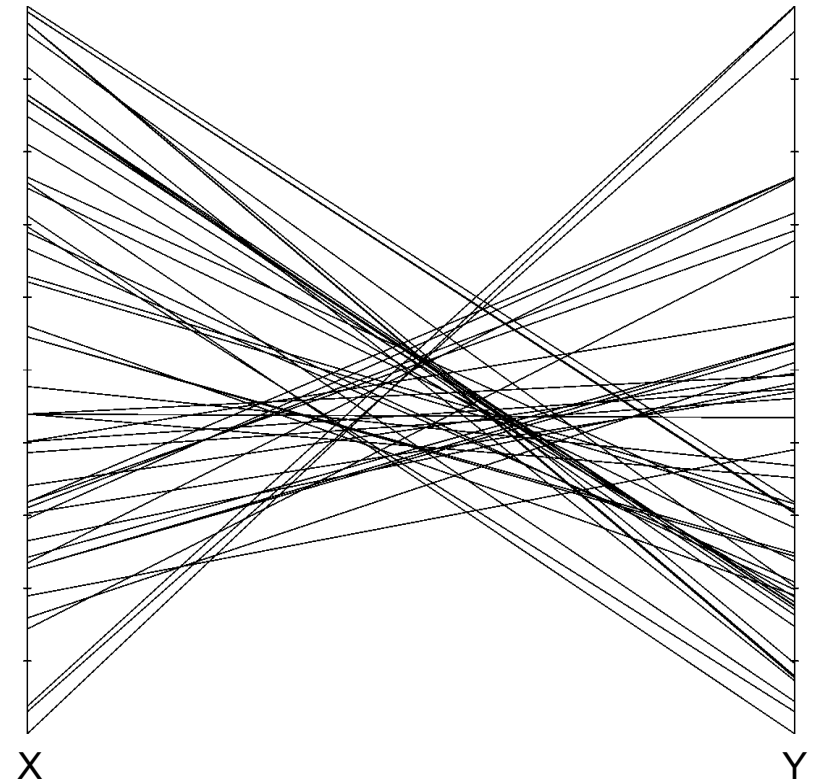
Reference: K. Wu, E. Otoo, and A. Shoshani, "Compressing bitmap indexes for faster search operations", ACM Transactions on Database Systems, vol 31, pp. 1-38, 2006

Introduction to Parallel Coordinates

2D Scatter-plot

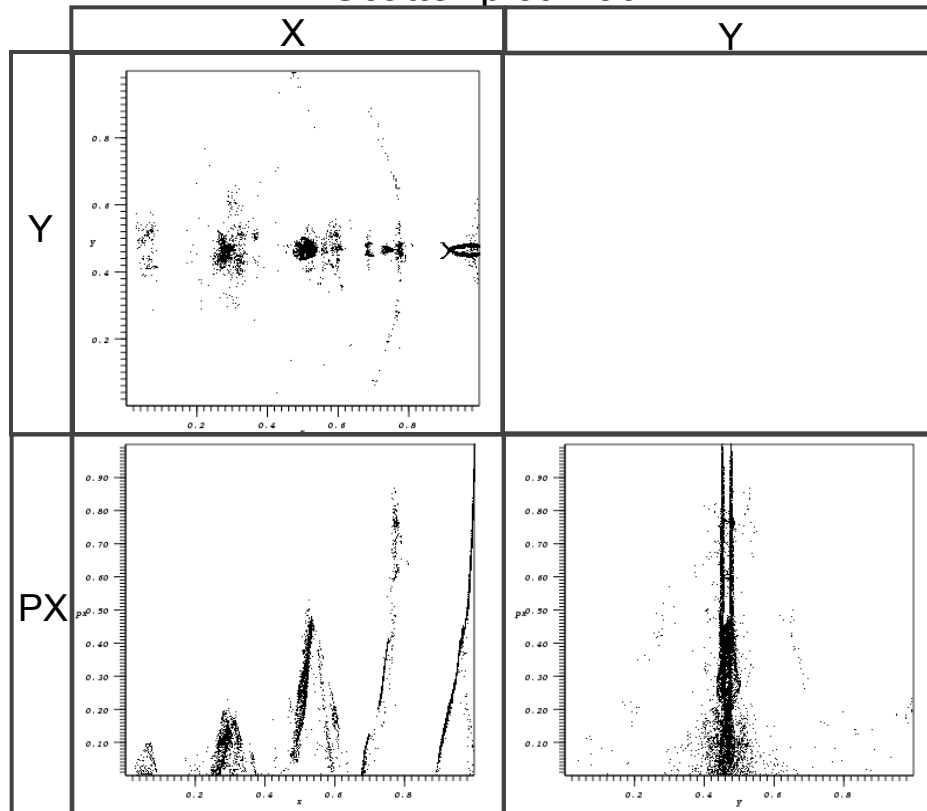


Parallel Coordinates

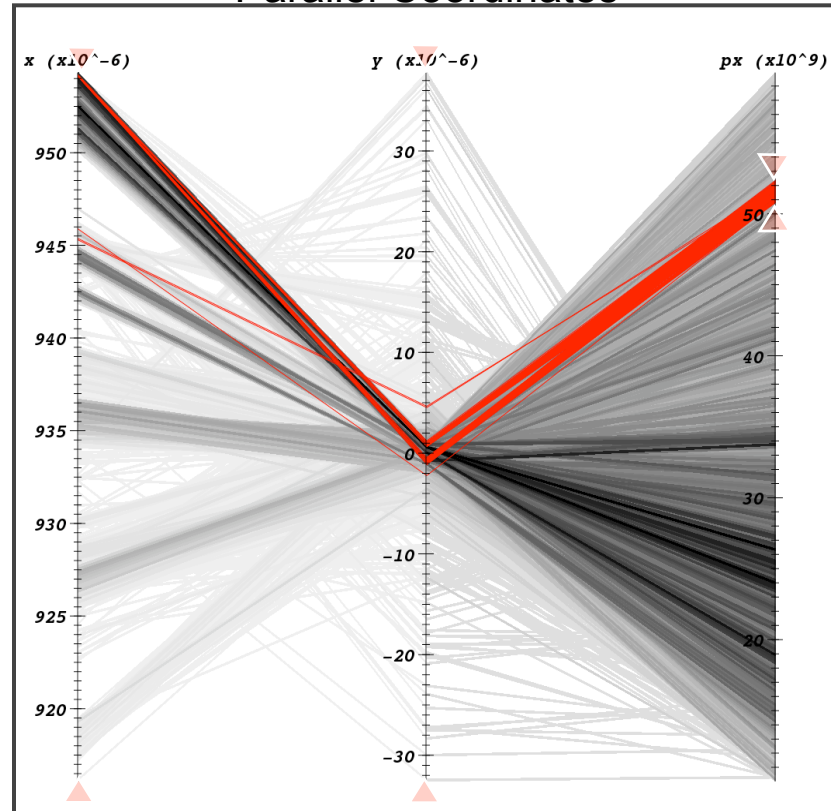


Introduction to Parallel Coordinates, cont.

Scatter-plot Matrix



Parallel Coordinates



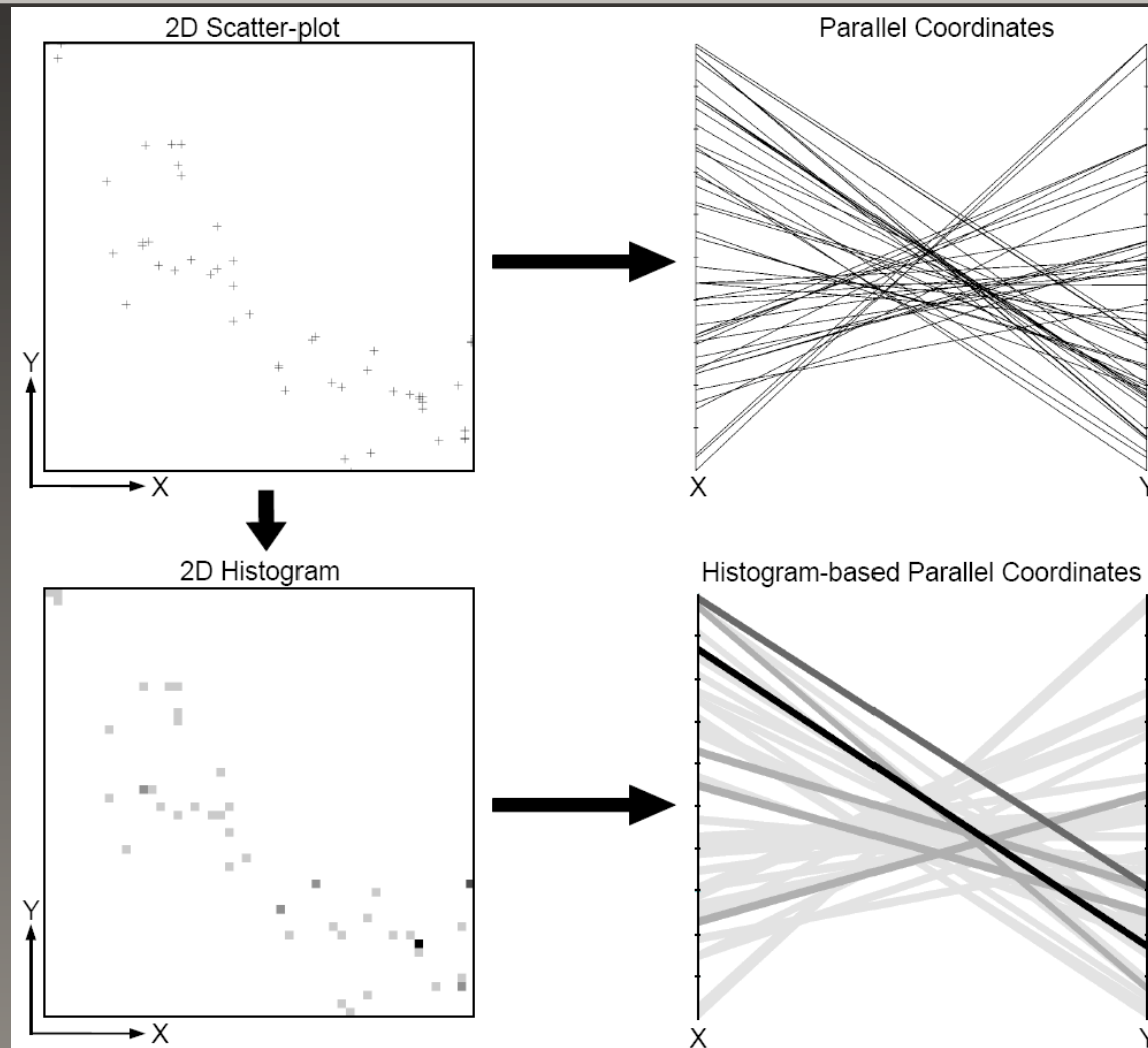
Advantages:

- Parallel display of many data dimensions
- Easy interface for data thresholding
- Immediate feedback during data selection

Disadvantages:

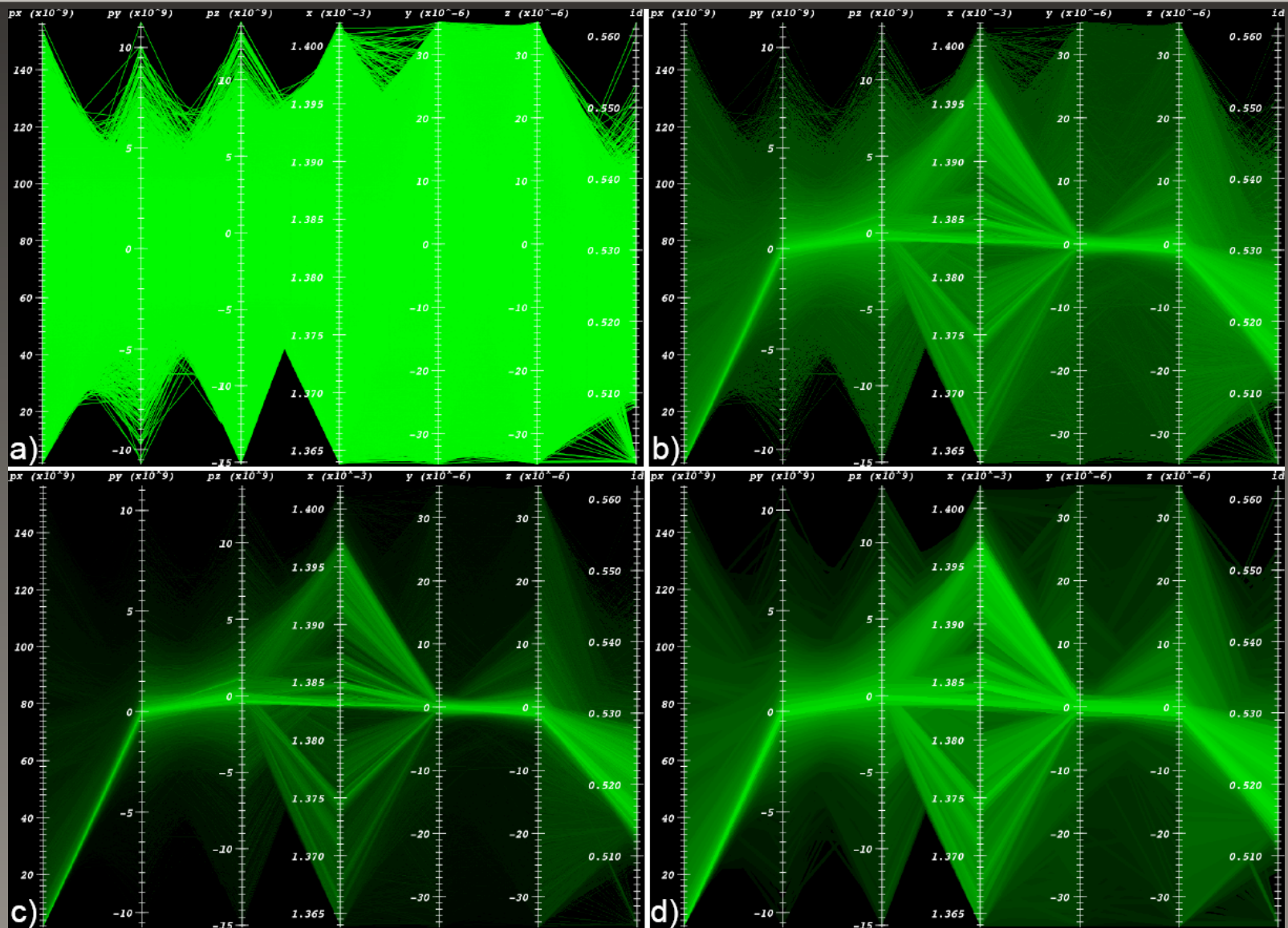
- Order dependent visualization
- Clutter, Occlusion
- Comp. complexity proportional to data size

Histogram-based Parallel Coordinates



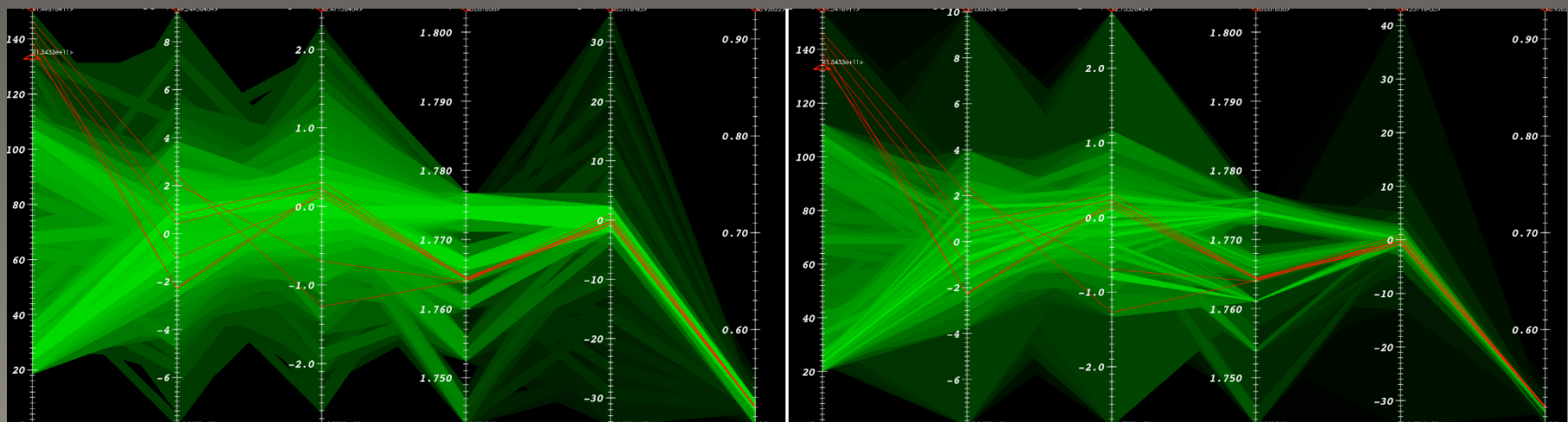
Reference: M. Novotny and H. Hauser, "Outlier-preserving focus+context visualization in parallel coordinates," *IEEE Transactions on Visualization and Computer Graphics*, vol. 12, no. 5, pp. 893-900. 2006.

Histogram-based Parallel Coordinates cont.



Histogram-based Parallel Coordinates, cont.

- Histograms computed on request:
 - Rendering of data subsets using histograms
 - Close zoom-ins and smooth drill-downs into the data
 - Rendering with arbitrary number of bins
- Support adaptively binned histograms:
 - More accurate representation in lower-level-of-detail views

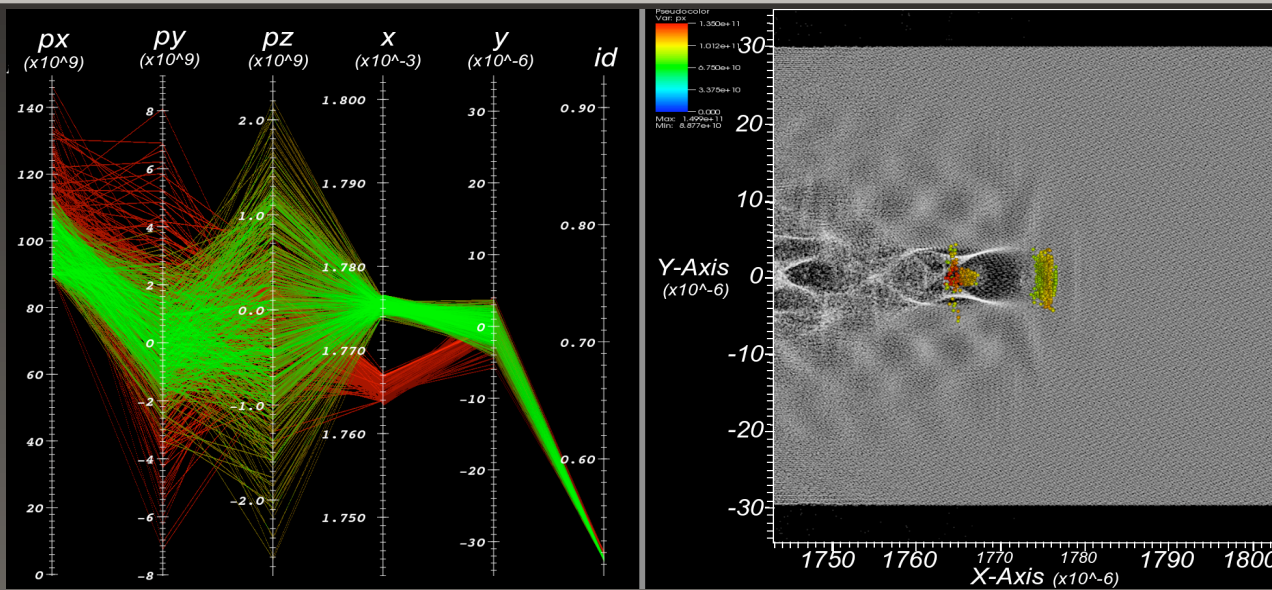


32x32 uniform binning

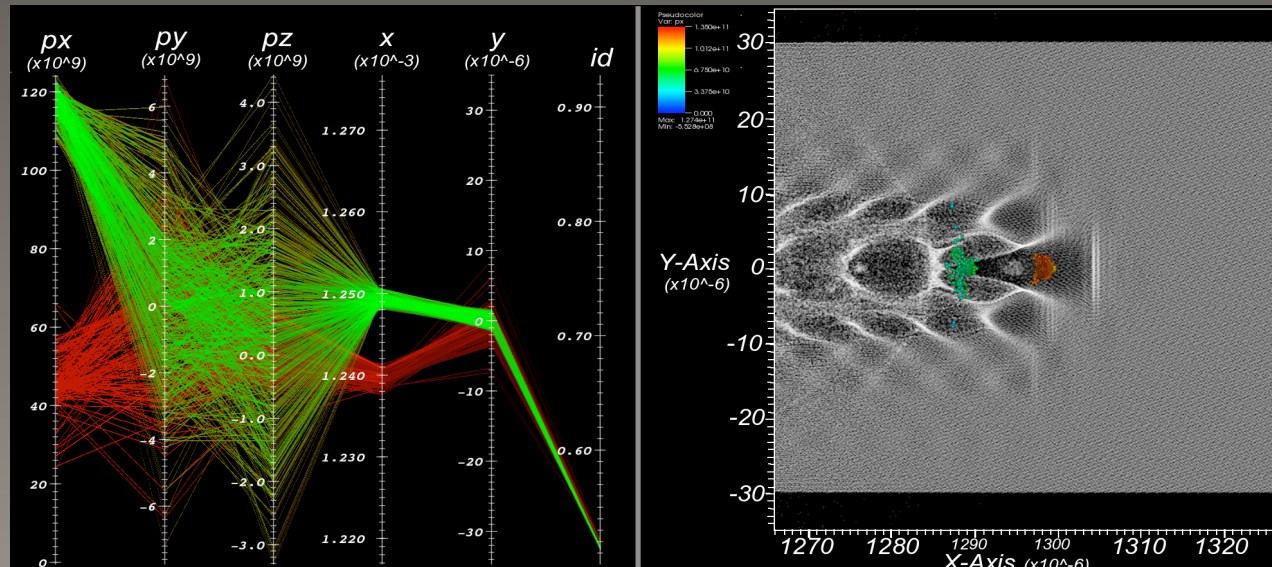
32x32 adaptive binning

Beam Selection and Assessment

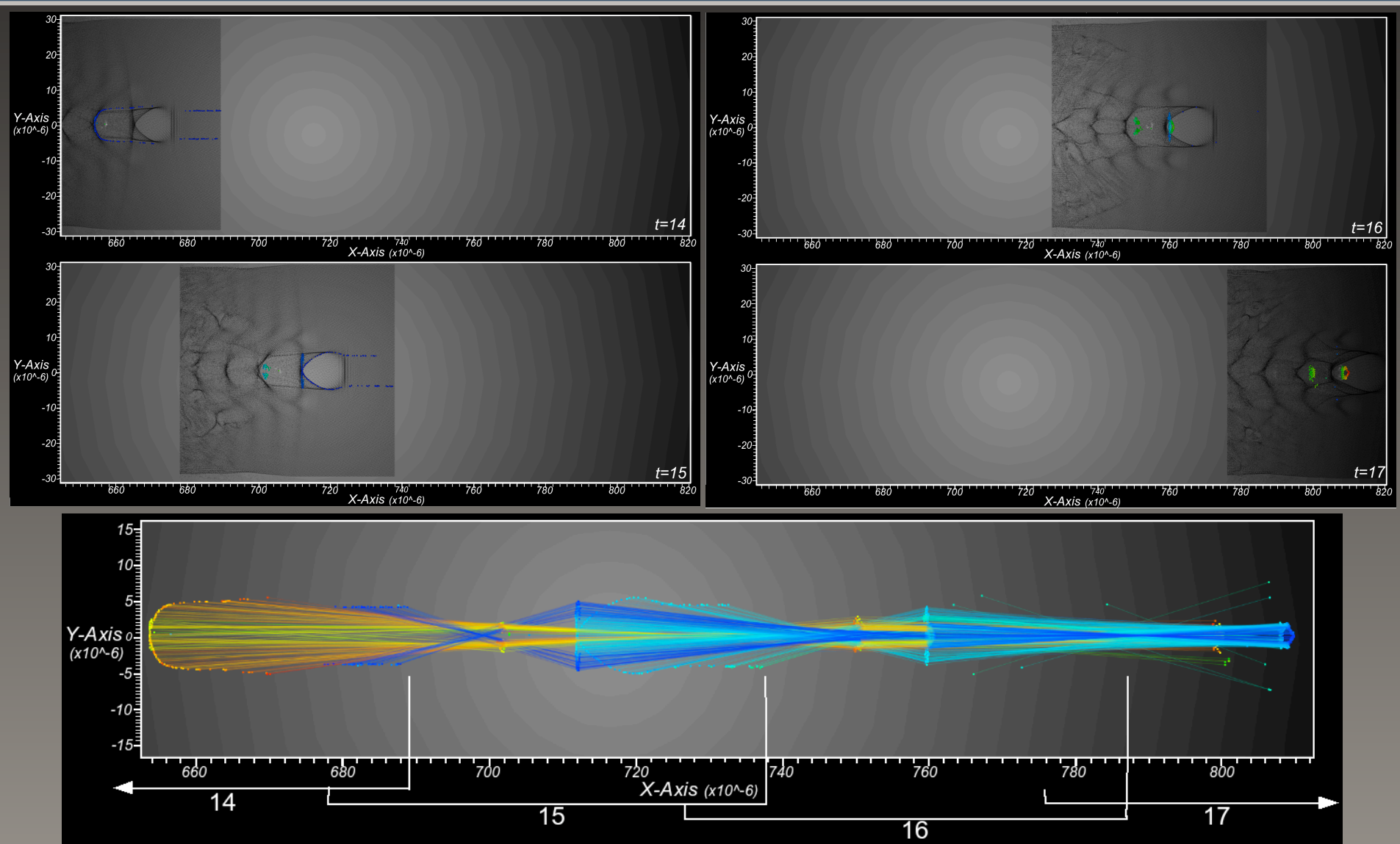
t=37



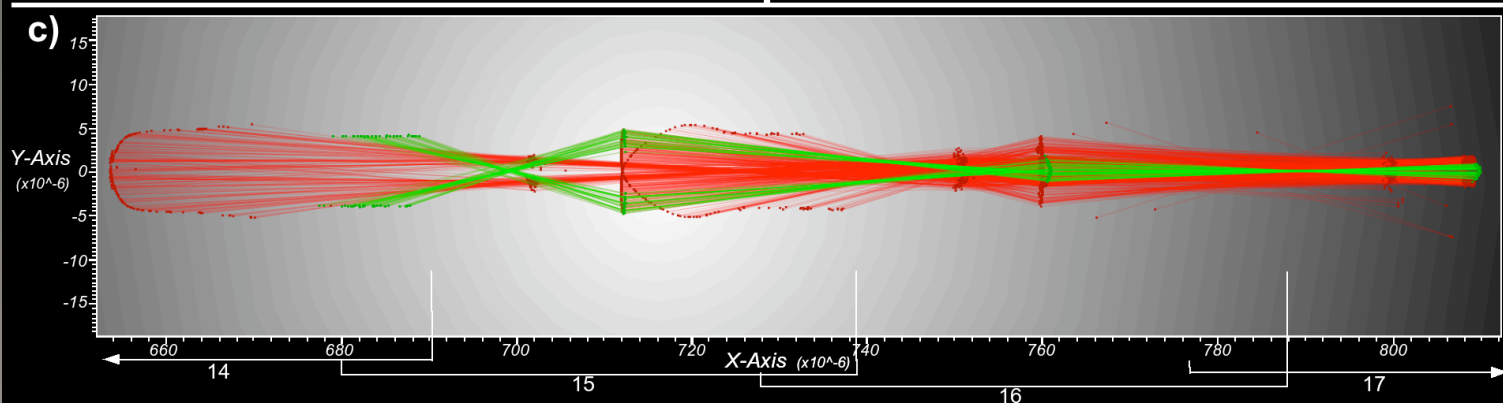
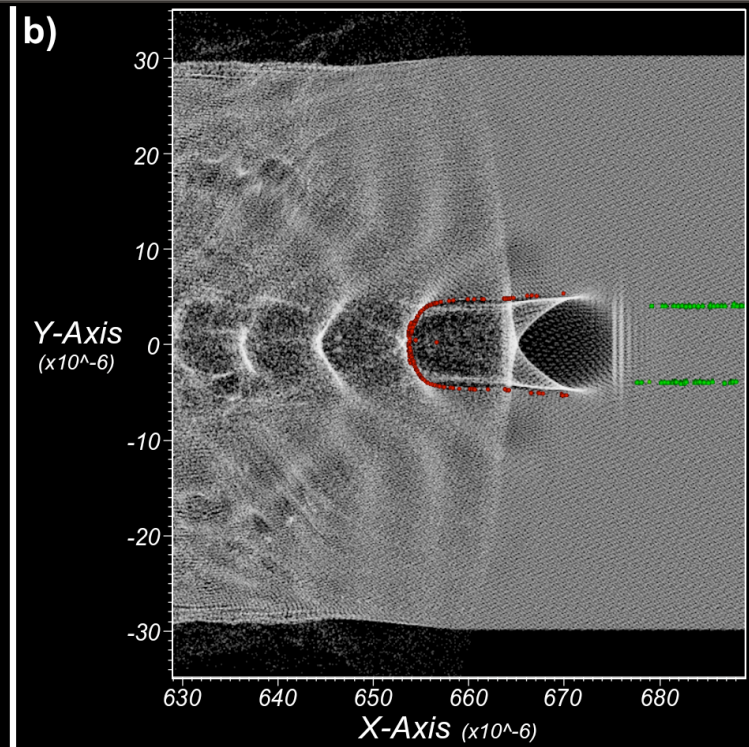
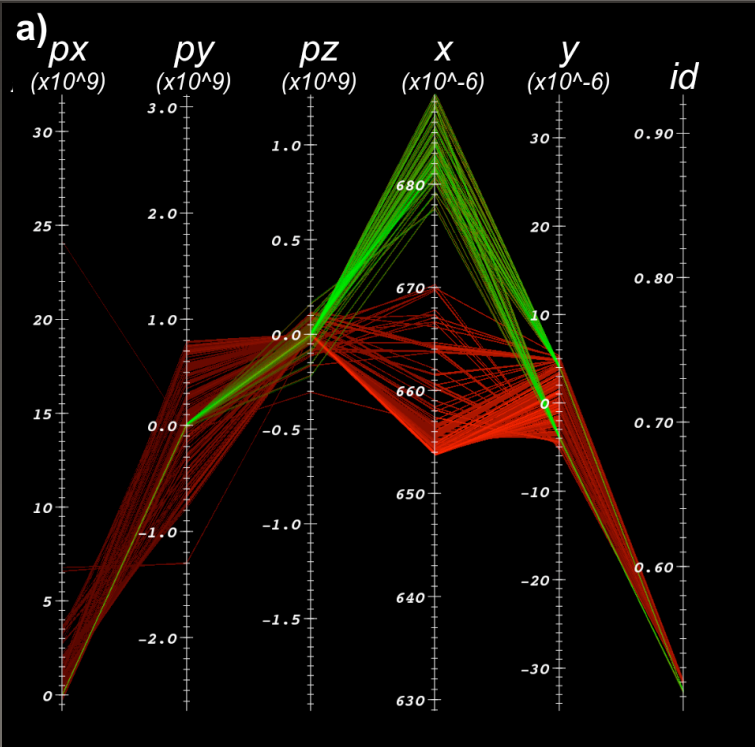
t=27



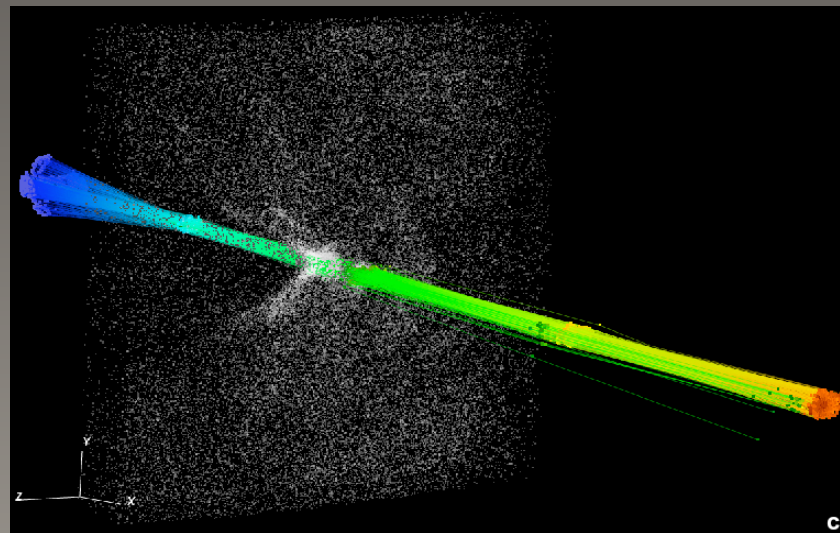
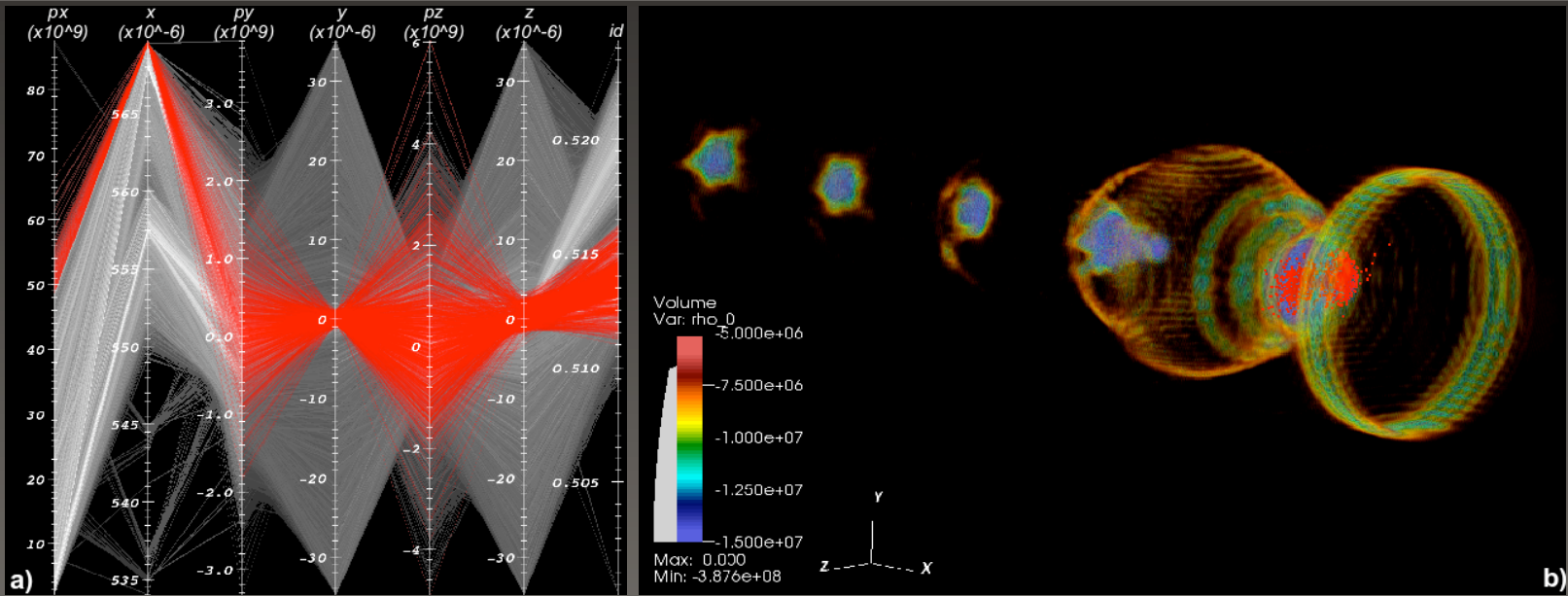
Beam Formation



Beam Refinement



3D Analysis Example



Laser Wakefield Acceleration Simulations – Conclusions and Future Work

- **Conclusions**

- Rapid knowledge discovery from large, complex, multivariate, time-varying data
- New approach for quickly generating histogram-based parallel coordinates
- Case study on how system can be used to analyze Laser Wakefield particle acceleration data effectively

- **Future Work**

- Distribute in public VisIt version (1.12)
- Support for more file formats, e.g., particles in Chombo files
- Explore parallelizing most expensive system parts
- Improve integration of field and particle data
- Couple with other traditional data analysis methods, e.g., clustering

Acknowledgements

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 - Director, Office of Advanced Scientific Computing Research, Office of Science, U.S. Department of Energy under Contract No. DE-AC02-05CH11231 through the Scientific Discovery through Advanced Computing (SciDAC) program's Visualization and Analytics Center for Enabling Technologies (VACET).
 - National Energy Research Scientific Computing Center, supported by Office of Science of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.
- **We would like to thank**
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 - Members of VACET
 - Members of LBNL Vis Group
 - Members of LBNL Center for Computational Sciences and Engineering

Questions?



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